

EFFECT OF POSITRON TEMPERATURE ON HIGH RELATIVISTIC ELECTRON-POSITRON-ION PLASMAS WITH NONTHERMAL ELECTRONS

R. Das and R. Sarma

Department of Mathematics, Arya Vidyapeeth College, Guwahati 781016, Assam, India

Abstract— In this model of relativistic electron-positron-ion plasma, only compressive solitons are found to exist. The amplitude becomes higher for smaller values of $\sigma = T_e / T_p$ (= electron to positron temperature ratio). The investigation further revealed that higher values of v_0 / c (v_0 is the initial ion streaming and c is the normalized velocity of light) gives higher amplitudes of compressive solitons.

Keywords: Relativistic effects, Positron temperature, Ion acoustic, Soliton.

1 INTRODUCTION

Waves in plasmas are an important phenomena interconnecting set of particles and fields which propagates periodically in a repeated fashion. A solitary wave is one propagates without any temporal evolution in shape and size and arises in many context like the elevation of the surface of water and the intensity of light in optical fibres. In the last few years, a considerable amount of interest has grown in the field of coherent nonlinear wave structures. Out of these nonlinear structures, ion acoustic solitons stands tall in modern plasma research. The first experimental observation of ion acoustic waves was made by Ikezi et al. [1]. Also many researchers [2, 3] have studied ion acoustic solitons using reductive perturbation method. A great deal of interest has risen recently in the study of nonlinear wave phenomena in electron-positron-ion plasma [4 – 9] which is mainly due to electron-positron-ion plasmas occur in many astrophysical environments such as active galactic nuclei [10], pulsar magnetosphere [11], polar regions of neutron stars [12], centre of our galaxy [13], the early universe [14,15] and solar atmosphere [16] and also produced in some laboratory environments [17 – 19]. Linear and nonlinear wave propagation in electron-positron and electron – positron – ion plasma has been studied using different models. For instance, Popel *et al.* [5] explored ion acoustic solitons in three component plasmas constituting electrons, positrons and singly charged ions giving a result that the presence of positron reduces ion acoustic amplitude. Moslem *et al.* [20] investigated the nonlinear two dimensional cylindrical acoustic excitations (solitons and double layers) in electron – positron – ion plasma having a composition of warm electrons and positive ions. However, most of these studies are focused on nonrelativistic plasmas, but when the particle velocities are comparable to the speed of light, relativistic effects may significantly modify the soliton behaviour [21 – 28]. Relativistic plasmas occur in a variety of situation like space plasmas [29], laser plasma interaction [30], plasma sheet boundary layer of earth's magnetosphere [31] and is also used in describing Van Allen Radiation belts [32]. Javidan and

Saadatmand [33] have studied the effect of high relativistic ions and nonthermal electrons in electron-ion-positron plasma system. They have obtained the maximum amplitude of the solitary wave and its width as functions of plasma parameters. Sabarian et al. [34] have investigated the propagation of large amplitude ion-acoustic solitary waves in fully relativistic plasma consisting of cold ions and ultra-relativistic hot electrons and positrons using Sagdeev pseudo potential method. Shah et al [35] have studied the propagation of ion acoustic solitary waves in a plasma system comprising of relativistic ions, kappa distributed electrons, and positrons. They reported that the increase in the relativistic streaming factor and positron and electron kappa parameters cause the soliton amplitude to thrive. However, the soliton amplitude diminishes as the positron concentration is increased in the system. Saeed et al. [36] have investigated about the Nonlinear Korteweg–de Vries equation for soliton propagation in relativistic electron-positron-ion plasma with thermal ions.

In this paper, we have investigated effect of electron to positron temperature ratio in high relativistic electron-positron-ion plasma.

2 BASIC EQUATIONS AND DERIVATION OF KdV EQUATION

We consider collisionless, unmagnetized plasma consisting of positive ions, thermal positrons and nonthermal electrons. The fluid equations of motion, governing the collisionless plasma in one dimension are:

$$\frac{\partial n_i}{\partial t} + \frac{\partial}{\partial x}(n_i v_i) = 0 \quad (1)$$

$$\left(\frac{\partial}{\partial t} + v_i \frac{\partial}{\partial x} \right) n_i + \frac{\partial \phi}{\partial x} = 0 \quad (2)$$

$$n_e = e^{\phi} \quad (3)$$

$$n_e = p e^{-\sigma\phi} \tag{4}$$

$$\frac{\partial^2 \phi}{\partial x^2} = n_e - n_i - n_p \tag{5}$$

where $\gamma = \left(1 - \frac{v_i^2}{c^2}\right)^{\frac{1}{2}} = 1 + \frac{v_i^2}{2c^2} + \frac{3v_i^4}{8c^4}$

where, i, p and e stand for positive ion, positron and electron respectively and $\sigma = T_e/T_p$ (=electron to positron temperature ratio).

We normalize densities of the plasma species by the unperturbed densities n_{e0} , time t by the inverse of the characteristic ion plasma frequency i.e., $\omega_{pi}^{-1} = \left(\frac{m_i}{4\pi n_{e0} e^2}\right)^{\frac{1}{2}}$, distance x by the electron Debye length $\lambda_{De} = \left(\frac{T_e}{4\pi n_{e0} e^2}\right)^{\frac{1}{2}}$, velocities by the ion-acoustic speed $C_s = \left(\frac{T_e}{m_i}\right)^{\frac{1}{2}}$, and the potential ϕ by $\frac{T_e}{e}$.

Introducing the stretched variables

$$\xi = \varepsilon^{\frac{1}{2}}(x - Ut), \tau = \varepsilon^{\frac{3}{2}}t \tag{6}$$

where ε is a small parameter which characterizes the strength of the nonlinearity and U is the phase velocity. Dependent variables are expanded as follows:

$$n_i = 1 - p + \varepsilon n_{i1} + \varepsilon^2 n_{i2} + \dots$$

$$n_e = 1 + \varepsilon n_{e1} + \varepsilon^2 n_{e2} + \dots$$

$$n_p = p + \varepsilon p_{p1} + \varepsilon^2 n_{p2} + \dots$$

$$v_i = v_0 + \varepsilon v_{i1} + \varepsilon^2 v_{i2} + \dots$$

$$\phi = \varepsilon \phi_1 + \varepsilon^2 \phi_2 + \dots$$

Using (6) and (7) in equations (1) - (5) and equating the coefficients of the lowest order perturbation in ε with the use of the boundary conditions $n_{e1} - n_{i1} - n_{p1} = 0, v_{i1} = 0, \phi_1 = 0$ at $|\eta| \rightarrow \infty$, we get

$$n_{i1} = \frac{1-p}{(U-v_0)^2 \beta} \phi_1, n_{e1} = \phi_1, n_{p1} = -\sigma p \phi_1, n_{e1} - n_{i1} - n_{p1} = 0 \tag{8}$$

From the last equation of (8), the expression for the phase velocity U can be written as

$$1 - \frac{1-p}{(U-v_0)^2 \beta} + \sigma p = 0 \tag{9}$$

This gives

$$U = v_0 \pm \sqrt{\frac{1-p}{\beta(1+\sigma p)}}, \text{ where } p = \frac{n_{p0}}{n_{e0}} \tag{10}$$

From the second order equations in ε , KdV equation can be obtained as

$$\frac{\partial \phi_1}{\partial \tau} + p' \phi_1 \frac{\partial \phi_1}{\partial \xi} + q' \frac{\partial^3 \phi_1}{\partial \xi^3} = 0 \tag{11}$$

where $p' = \frac{A}{B}$ and $q' = \frac{1}{B}$

$$\text{with } A = \frac{3(1-p)}{(U-v_0)^4 \beta^2} - \frac{2\beta_0(1-p)}{(U-v_0)^4 \beta^3} - (1-\sigma^2 p)$$

$$\text{and } B = \frac{2(1-p)}{(U-v_0)^3 \beta}$$

3 SOLITARY WAVE SOLUTION

The solution of the KdV equation (11) is obtained as

$$\phi_1 = \frac{3V}{p} \operatorname{sech}^2 \left(\frac{1}{2} \sqrt{\frac{V}{q}} \chi \right) \tag{12}$$

where V represents the speed of the ion acoustic soliton

with amplitude $\phi_0 = \frac{3V}{p}$ and width $\Delta = 2\sqrt{\frac{q}{V}}$. In deducing

the solution (12), a new variable $\chi = \eta - V\tau$ is introduced and

following boundary conditions are used $\phi_1 = \frac{\partial\phi_1}{\partial\eta} = \frac{\partial^2\phi_1}{\partial\eta^2} = 0$

at $|\eta| \rightarrow \pm\infty$.

4 DISCUSSION

In this model of plasma, we have investigated the ion acoustic solitary waves in high relativistic electron-positron-ion plasma considering electron to positron temperature ratio σ , subject to the condition $0 < p \leq 1$. The plasma parameters such as the relativistic factor v_0/c , temperature ratio σ and positron concentration p are found to play a significant role in the formation of ion acoustic solitons. The amplitude of compressive KdV soliton [Fig. 1(a)] increases as v_0/c increases for fixed values of $V = 0.10$, $p = 0.50$ and for different values of $\sigma = 0.05, 0.10, 0.15$. From figure it is revealed that for higher temperature the amplitude becomes smaller. On the other hand, the width [Fig. 1(b)] of the compressive soliton decreases with v_0/c for the same set of parametric values. The amplitudes [Fig. 2(a)] of the compressive solitons showing linear decrease with σ for fixed $V = 0.20$, $v_0/c = 0.30$ and for different values of $p = 0.10(1), 0.20(2), 0.30(3)$. The corresponding widths [Fig. 2(b)] of the fast compressive solitons also decrease as figure 2(a). The amplitudes (ϕ_0) [Fig. 3 (a)] of the compressive KdV solitons are seen to decrease sharply but linearly with p for fixed $V = 0.20$, $v_0/c = 0.30$ and for different values of $\sigma = 0.10, 0.30, 0.50$. But the corresponding widths (Δ) [Fig. 3(b)] decrease concavely with p . Figure 4 shows that the soliton height and width decreases as positron concentration p increases. This situation indicates that the soliton energy decreases with an increasing p .

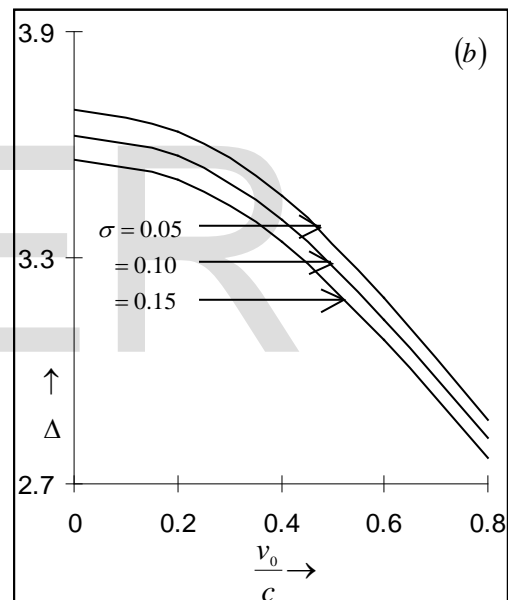
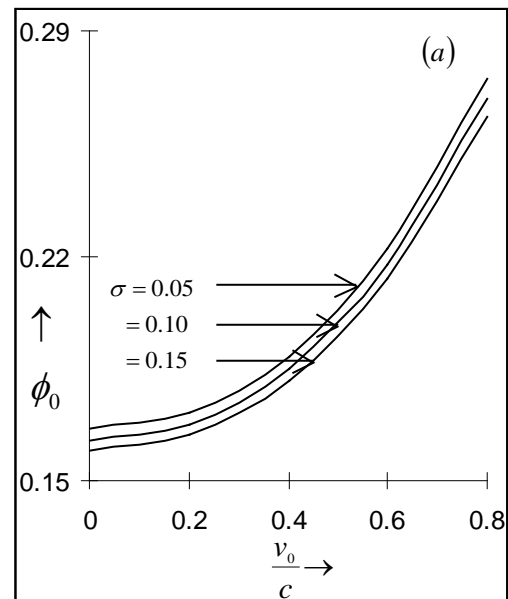


FIG.1. Amplitudes (a) and widths (b) of higher order relativistic compressive KdV solitons versus $\frac{v_0}{c}$ for fixed $V = 0.10$, $p = 0.50$ and for different values of $\sigma = 0.05, 0.10, 0.15$.

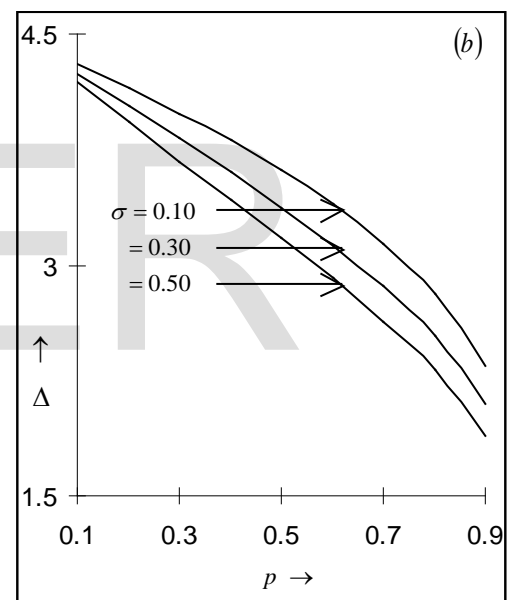
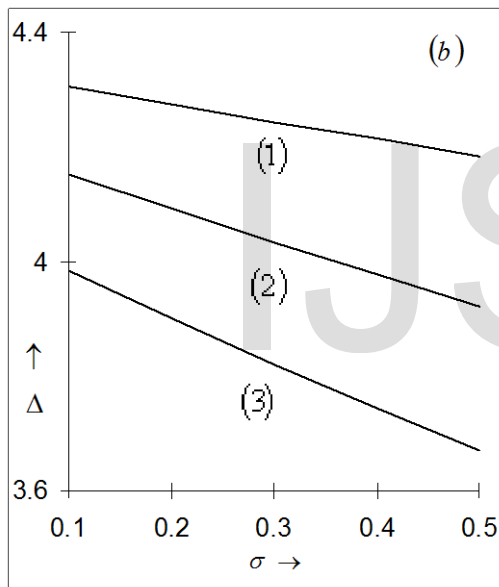
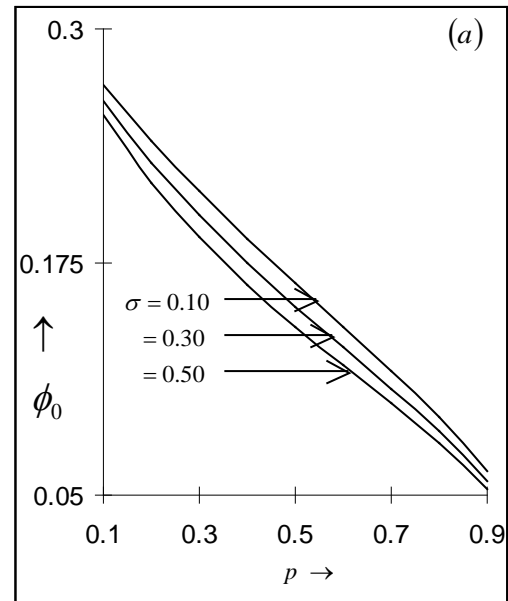
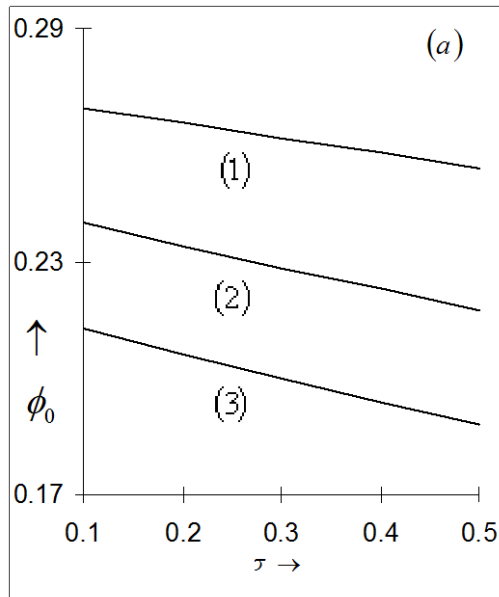


FIG.2. Amplitudes (a) and widths (b) of higher order relativistic compressive KdV solitons versus temperature ratio σ for fixed $V = 0.20$, $\frac{v_0}{c} = 0.30$ and for different values of $p = 0.10(1), 0.20(2), 0.30(3)$.

FIG.3. Amplitudes (a) and widths (b) of higher order relativistic compressive KdV solitons versus positron concentration p for fixed $V = 0.20$, $\frac{v_0}{c} = 0.30$ and for different values of $\sigma = 0.10, 0.30, 0.50$.

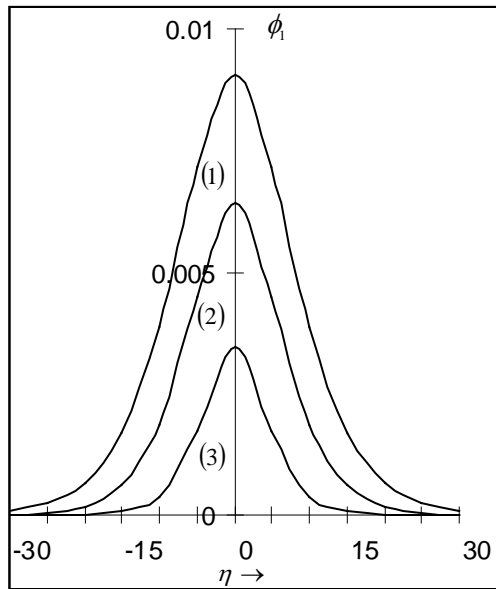


FIG.4. Plot of the amplitude of the compressive KdV solitons with $V = 0.0075$, $\sigma = 1$ and $v_0/c = 0.30$ for different values of $p = 0.50(1), 0.70(2), 0.90(3)$.

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REFERENCES

- [1] H. Ikezi, R. J. Taylor, and D. R. Baker, Formation and Interaction of Ion-Acoustic Solitons, *Phys. Rev. Lett.* 25, 11 – 14 (1970).
- [2] R. Bharuthram and P. K. Shukla, Large amplitude ion-acoustic double layers in a double Maxwellian electron plasma, *Phys. Fluids* 29, 3214 – 3218 (1986).
- [3] L. L. Yadav and S. R. Sharma, Obliquely propagating ion-acoustic double layers in a multicomponent magnetized plasma, *Phys. Scr.* 43, 106 – 110 (1991).
- [4] P. K. Shukla, A. A. Mamun and L. Stenflo, Vortices in a Strongly Magnetized Electron-Positron-Ion Plasma, *Phys. Scr.* 68, 295 – 297 (2003).
- [5] S. I. Popel, S. V. Vladimirov, and P. K. Shukla, Ion-acoustic solitons in electron-positron-ion plasmas, *Phys. Plasmas* 2, 716 – 719 (1995).
- [6] Y. N. Nejoh, The effect of the ion temperature on large amplitude ion-acoustic waves in an electron-positron-ion plasma, *Phys. Plasmas* 3, 1447 – 1451 (1996).
- [7] M. Mishra, A. K. Arora and R.S. Chhabra, Ion-acoustic compressive and rarefactive double layers in a warm multi-component plasma with negative ions, *Phys. Rev. E*, 66, 46402 (2002).
- [8] S. Ghosh, R. Bharuthram, Ion acoustic solitons and double layers in electron-positron-ion plasmas with dust particulates, *Astrophys. Space Sci* 314, 121 – 127 (2008).
- [9] H.R. Pakzad, Ion acoustic solitary waves in plasma with nonthermal electron, positron and warm ion, *Astrophys. Space Sci.* 323 345 – 350 (2009).
- [10] H.R. Miller, P.J. Witta, *Active Galactic Nuclei*, P.202 Springer Berlin (1978).
- [11] F.C. Michel, Theory of pulsar magnetospheres, *Rev Mod Phys.* 54, 1(1982).
- [12] F.C. Michell, *Theory of Neutron Star Magnetosphere*, Chicago University Press Chicago (1991)
- [13] M.I. Barnes, *Positron electron pairs in Astrophysics*, New York : American Institute of Physics (1983).
- [14] W.K. Misner, S. Thorns, J.A. Wheeler, *Gravitation*, P 763 Freeman San Francisco (1973).
- [15] M.J. Rees, G.W. Gibbons, S.W. Hawking, S. Siklaseds, *The*

Early Universe, Cambridge University Press, Cambridge (1995).

(1983).

[16] Tandberg, E Hansen, A.G. Emslie The physics of solar Flares, Cambridge University Press Cambridge, 124 (1988).

[17] C. M. Surko, M. Leventhal and A. Passner, Positron Plasma in the Laboratory, Phys. Rev. Lett. 62, 901 – 904 (1989).

[18] H.Bochmer, M.Adams, N.Rynn , Positron trapping in a magnetic mirror configuration, Phys. Plasmas 2, 4369 – 4371 (1995).

[19] E.P.Liang , S.C. Wilks and M.Tabak, Pair Production by Ultraintense Lasers ,Phys. Rev. Lett, 81, 4887 – 4890 (1998).

[20] W.M.Moslem, I.Kourakis, P.K. Shukla, R.Schlickeiser, Phys. Plasmas, 14, 102901 (2007).

[21] G. C. Das and S. N. Paul, Ion-acoustic solitary waves in relativistic plasmas, Phys. Fluids, 28, 823 - 825 (1985).

[22] Y. Nejoh, The effect of the ion temperature on the ion acoustic solitary waves in a collisionless relativistic plasma, J. Plasma Phys. 37, 487 - 495 (1987).

[23] G. C. Das, B. Karmakar and S. N. Paul, Propagation of solitary waves in relativistic plasmas, IEEE Trans. Plasma Sci. 16, 22 - 26 (1988).

[24] P. Chatterjee and R. Roychoudhury, Effect of ion temperature on large - amplitude ion-acoustic solitary waves in relativistic plasma, Phys. Plasmas 1, 2148 – 2153 (1994).

[25] S. K. EL-Labany and S. M. Shaaban, Contribution of higher order nonlinearity to nonlinear ion-acoustic waves in a weakly relativistic warm plasma, J. Plasma Phys. 53, 245 – 252

[26] K. Singh, V. Kumar, H.K. Malik, Electron inertia effect on small amplitude solitons in a weakly relativistic two-uid plasma, Phys. Plasmas 12 , 052103 (2005).

[27] T. S. Gill, H. Kaur, N.S.J. Saini, A study of ion-acoustic solitons and double layers in a multispecies collisionless weakly relativistic plasma, J. Plasma Phys. 71, 23 - 34 (2005).

[28]T. S. Gill, A. Singh, H. Kaur, N. S. Saini and P. Bala, Ion-acoustic solitons in weakly relativistic plasma containing electron-positron and ion, Phys. Letters A 361, 364 – 367 (2007).

[29] J. Arons, Some problems of pulsar physics or I'm madly in love with electricity, Space Sci. Rev. 24, 437 - 510 (1979).

[30] C. Grabbe, Wave propagation effects of broadband electrostatic noise in the magnetotail, J. Geophys. Res. 94, 17299 – 17304(1989)

[31] J I Vette Summary of particle population in the magnetosphere. Particles and Fields in the Magnetosphere (Dordrecht: Reidel)p 305 (1970)

[32] H Ikezi, Experiments on ion-acoustic solitary waves, Phys. Fluids 16 1668 – 1675 (1973)

[33] Kurosh Javidan and Danial Saadatmand, Effect of high relativistic ions on ion acoustic solitons in electron-ion-positron plasmas with nonthermal electrons and thermal positrons, Astrophys Space Sci 333, 471–475 (2011)

[34] E. Saberian, A. Esfandyari-Kalejahi and M. Akbari-Moghanjoughi, Propagation of ion-acoustic solitary waves in a relativistic electron-positron-ion plasma, Canadian Journal of

[35] Asif Shah, S. Mahmood, and Q. Haque, Propagation of solitary waves in relativistic electron-positron-ion plasmas with kappa distributed electrons and positrons, Phys. Plasmas 18, 114501 (2011)

[36] R. Saeed, Asif Shah and Muhammad Noaman-ul-Haq, Nonlinear Korteweg–de Vries equation for soliton propagation in relativistic electron-positron-ion plasma with thermal ions, Phys, Plasmas 17, 102301(2010).

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